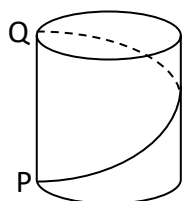


## § Geodesics

§ 01 Geodesic equation :  $\ddot{x}^i + \Gamma_{jk}^i \dot{x}^j \dot{x}^k = 0$

一曲面上連接兩點 有最短距離的曲線，稱為測地線 (geodesic)。

這樣的觀念在高等數學很重要，在初等數學 我們找到一些例子。



左圖一隻螞蟻從一圓柱體的糖罐底部 P 點沿著柱體側面旋轉  
往上爬到 Q 點，若柱體高  $\overline{PQ}=8$  公分，底面圓周長 6 公分，  
求螞蟻爬行的最短距離\_\_公分。

(剪開 攤開 10 公分)

這個例子也說明測地線不一定是最短路徑，螞蟻直接從 P 往上走到 Q，顯然距離比較近。

在相對論中，等價(效)原理(equivalence principle)指出重力場決定伽利略時空的曲線，這些曲線是對稱聯絡(Cartan connection)的測地線。

在廣義相對論中，一個自由粒子走的路徑是測地線。

An important property of geodesic in a spacetime with Lorentzian metric is that the character (timelike/null/spacelike) of the geodesic relative to a metric compatible connection, never changes。

This is simple because parallel transport preserves inner product, and the character is determined by the inner product of the tangent vector itself。

§ 02 Viewpoint of classical mechanism

In Lagrangian mechanism,  $C = \{c | c : [a, b] \rightarrow M, c(a) = p, c(b) = q\}$

A Lagrangian function L on M,  $L : TM \rightarrow R$

The action determined by L on C,  $S : C \rightarrow R$ ,  $S(c) := \int_a^b L(\dot{c}(t)) dt$

$\gamma : (-\varepsilon, \varepsilon) \rightarrow C$  a variation of c

such that  $\gamma(0) = c$  and  $\gamma(s)(t) : (-\varepsilon, \varepsilon) \times [a, b] \rightarrow M$  is differentiable

$\tilde{\gamma}(s, t) := \gamma(s)(t)$  . If  $\left. \frac{d}{ds} S(\gamma(s)) \right|_{s=0} = 0$  for  $\forall \gamma \in C$  , then curve  $c$  is called the critical point of the action  $S$  . (Least action principle)

Theorem

$c \in C$  is the critical point of  $S$  which determined by the Lagrangian  $L \Leftrightarrow c$  satisfies Euler-Lagrange equations .

$\gamma : [a, b] \rightarrow M$  is a smooth curve on Riemann manifold

$L(\gamma) := \int_a^b \left| \frac{d\gamma}{dt}(t) \right| dt$  is the length of  $\gamma$

$E(\gamma) := \frac{1}{2} \int_a^b \left| \frac{d\gamma}{dt}(t) \right|^2 dt = \frac{1}{2} \int_a^b (g_{ij} \dot{x}^i \dot{x}^j) dt$  is the energy of  $\gamma$

( In physics,  $E(\gamma)$  is usually called the “action of  $\gamma$ ” , where  $\gamma$  is considered as the orbit of a mass point . )

Represent by coordinates  $(x^1(\gamma(t)), x^2(\gamma(t)), \dots, x^m(\gamma(t)))$

$\dot{x}^i(t) := \frac{d}{dt}(x^i(\gamma(t)))$  then

$L(\gamma) = \int_a^b \sqrt{g_{ij}(x(\gamma(t))) \dot{x}^i(t) \dot{x}^j(t)} dt$  ,  $E(\gamma) = \frac{1}{2} \int_a^b g_{ij}(x(\gamma(t))) \dot{x}^i(t) \dot{x}^j(t) dt$

Then Euler-Lagrange equation of energy  $E$  is

$\ddot{x}^i(t) + \Gamma_{jk}^i(x(t)) \dot{x}^j(t) \dot{x}^k(t) = 0$  ,  $i=1,2,3,\dots$  i.e. geodesic equation

Proof

The Euler-Lagrange equation of a functional  $I(x) = \int_a^b f(t, x(t), \dot{x}(t)) dt$

Is given by  $\frac{d}{dt} \frac{\partial f}{\partial \dot{x}^i} - \frac{\partial f}{\partial x^i} = 0$

$E(\gamma) = \frac{1}{2} \int g_{jk}(x(t)) \dot{x}^j \dot{x}^k dt$  (  $g_{ij} \dot{x}^i \dot{x}^j = L(t, x(t), \dot{x}(t))$  ) , and the Euler-Lagrange

equation is  $\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{x}^i} \right) - \frac{\partial L}{\partial x^i} = 0$  , then  $\frac{d}{dt} (g_{ik} \dot{x}^k + g_{ji} \dot{x}^j) - g_{jk,i} \dot{x}^j \dot{x}^k = 0$

For  $i=1,2,\dots,m$  therefor

$$g_{ik}\ddot{x}^k + g_{ji}\ddot{x}^j + g_{ik,\ell}\dot{x}^\ell\dot{x}^k + g_{ji,\ell}\dot{x}^\ell\dot{x}^j - g_{jk,i}\dot{x}^j\dot{x}^k = 0.$$

Renaming some indices and using the symmetry  $g_{ik} = g_{ki}$ , we get

$$2g_{\ell m}\ddot{x}^m + (g_{\ell k,j} + g_{j\ell,k} - g_{jk,\ell})\dot{x}^j\dot{x}^k = 0, \quad \ell = 1, \dots, d, \quad (1.4.16)$$

and from this

$$g^{i\ell}g_{\ell m}\ddot{x}^m + \frac{1}{2}g^{i\ell}(g_{\ell k,j} + g_{j\ell,k} - g_{jk,\ell})\dot{x}^j\dot{x}^k = 0, \quad i = 1, \dots, d.$$

Because of

$$g^{i\ell}g_{\ell m} = \delta_{im}, \quad \text{and thus } g^{i\ell}g_{\ell m}\ddot{x}^m = \ddot{x}^i,$$

We get  $\ddot{x}^i(t) + \Gamma_{jk}^i(x(t))\dot{x}^j(t)\dot{x}^k(t) = 0$

Q.E.D. [Geometric analysis] p.23

Thus, geodesics are the critical points of the energy function.

Geodesics are paths of minimal energy.

§ 03

$$\frac{d^2u^i}{ds^2} + \Gamma_{jk}^i \frac{du^j}{ds} \frac{du^k}{ds} = 0 \cdots \text{equation of a geodesic (1)}$$

$$\dot{V}^i + \sum_{j,k} \Gamma_{jk}^i \dot{x}^j V^k = 0 \quad Y \text{ parallel transported along } X \quad (2)$$

$$\text{Covariant derivative of } Y \text{ along } X, \quad \nabla_X Y = \sum_i (XY^i + \sum_{j,k} \Gamma_{jk}^i X^j Y^k) \frac{\partial}{\partial x^i} \quad (3)$$

(1) (2) (3) are mutually determined.

§ 04 A geodesic is a curve along which the tangent vector is parallel transported.

$$\nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j} := \Gamma_{jk}^i \frac{\partial}{\partial x^k} \quad \text{then}$$

$$\nabla_{\frac{\partial}{\partial x^i}} (w^j \frac{\partial}{\partial x^j}) = w^j \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j} + \frac{\partial w^j}{\partial x^i} \frac{\partial}{\partial x^j} = w^j \Gamma_{jk}^i \frac{\partial}{\partial x^k} + \frac{\partial w^j}{\partial x^i} \frac{\partial}{\partial x^j}$$

$$\text{And } \nabla_{v^i \frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j} = v^i \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j}$$

$$\text{For a curve } \mathbf{x}(t), \text{ with } \frac{d}{dt} = \frac{dx^i}{dt} \frac{\partial}{\partial x^i}$$

$$\nabla_{\frac{d}{dt}} \frac{dx}{dt} = \left( \frac{d^2x^i}{dt^2} + \Gamma_{jk}^i \frac{dx^j}{dt} \frac{dx^k}{dt} \right) \frac{\partial}{\partial x^i}, \quad \frac{d^2x^i}{dt^2} + \Gamma_{jk}^i \frac{dx^j}{dt} \frac{dx^k}{dt} = 0 \quad \text{is the geodesic equation}$$

That is  $\nabla_x X = 0$

§ 05 covariant derivative 平行移動 協變微分

$$\nabla_\mu V^\nu = \partial_\mu V^\nu + \Gamma_{\mu\lambda}^\nu V^\lambda$$

Parallel transport of the tensor T along the path  $x^\mu(\lambda)$ ,  $\frac{D}{d\lambda} T = 0$

$$\frac{d}{d\lambda} V^\mu + \Gamma_{\sigma\rho}^\mu \frac{dx^\sigma}{d\lambda} V^\rho = 0 \text{ for a vector } V^\mu$$

$$\frac{D}{d\lambda} \frac{dx^\mu}{d\lambda} = 0 \text{ or}$$

$$\frac{d^2 x^\mu}{d\lambda^2} + \Gamma_{\sigma\rho}^\mu \frac{dx^\sigma}{d\lambda} \frac{dx^\rho}{d\lambda} = 0$$

§ 06 Minimal-coupling principle 最小耦合原理

就是說 粒子在不受力的情形下

(1)寫出平直(歐氏)空間的運動方程式(直線)

(2)把加速度中的微分改成協變微分

(3)得到彎曲空間中的運動方程式(即 測地線方程式)

$$\frac{d^2 x^\mu}{d\lambda^2} = 0 \text{ for a free particle}$$

$$\rightarrow \frac{dx^\nu}{d\lambda} \partial_\nu \frac{dx^\mu}{d\lambda} \rightarrow \frac{dx^\nu}{d\lambda} \nabla_\nu \frac{dx^\mu}{d\lambda} = \frac{d^2 x^\mu}{d\lambda^2} + \Gamma_{\sigma\rho}^\mu \frac{dx^\sigma}{d\lambda} \frac{dx^\rho}{d\lambda}$$

§ 07 method of variation 變分法

$$\delta \int_a^b \sqrt{g_{ij} \dot{x}^i \dot{x}^j} dt = 0 \quad \text{Yves Talpaert p.30}$$

[Differential Geometry with Applications to Mechanics and Physics]

$$\text{Euler-Lagrange equation } \frac{d}{dt} \left( \frac{\partial F}{\partial \dot{x}^i} \right) - \frac{\partial F}{\partial x^i} = 0, \text{ with } F = \sqrt{g_{ij} \dot{x}^i \dot{x}^j}$$

We establish the geodesic equations from the calculus of variations .

On a manifold , a geodesic joining two points  $x_a, x_b$  is a curve such that

$$\delta \int_{s_a}^{s_b} ds = \delta \int_a^b \sqrt{g_{ij} \frac{dx^i}{dt} \frac{dx^j}{dt}} dt = 0$$

The calculus of variations means it is a curve “satisfying” the Euler

$$\text{equation } \frac{d}{dt} \left( \frac{g_{ij} \dot{x}^j}{F} \right) - \frac{1}{2F} \partial_i (g_{jk}) \dot{x}^j \dot{x}^k = 0$$

Substituting the curvilinear parameter  $s$  for  $t$ , the geodesic equations are

$$\frac{d}{dt} \left( g_{ij} \frac{dx^j}{dt} \right) - \frac{1}{2ds} \partial_i (g_{jk}) \frac{dx^j}{dt} \frac{dx^k}{dt} = 0$$

$$g_{ij} \frac{d^2 x^j}{ds^2} + \partial_k g_{ij} \frac{dx^k}{ds} \frac{dx^j}{ds} - \frac{1}{2} \partial_i g_{jk} \frac{dx^j}{ds} \frac{dx^k}{ds} = 0$$

$$g_{ij} \frac{d^2 x^j}{ds^2} + \frac{1}{2} (\partial_k g_{ij} \frac{dx^k}{ds} \frac{dx^j}{ds} + \partial_j g_{ik} \frac{dx^j}{ds} \frac{dx^k}{ds}) - \frac{1}{2} \partial_i g_{jk} \frac{dx^j}{ds} \frac{dx^k}{ds} = 0$$

By multiplying by  $g^{ri}$ , we obtain the geodesic equations

$$\frac{d^2 x^r}{ds^2} + \Gamma_{jk}^r \frac{dx^j}{ds} \frac{dx^k}{ds} = 0, \text{ where } s \text{ is proportional to } t.$$

看起來與§ 02 相同。

§ 08 heat flow method 從熱力學 heat flow 方程看測地線

Heat equation on  $\mathbf{R}$

$$(1) u_t - ku_{xx} = 0$$

$$(2) \begin{cases} u_t - ku_{xx} = f(x, t), t > 0 \\ u|_{t=0} = g(x) \end{cases}$$

Consider  $u_t^i = u_{ss}^i + \Gamma_{jk}^i u_s^j u_s^k, u(s, 0) = \gamma(s)$  for  $s \in S^1$  then

1.  $u_t^i - u_{ss}^i = f$  is a heat equation
2.  $u_{ss}^i + \Gamma_{jk}^i u_s^j u_s^k = 0$  is a geodesic equation

c.f.

1. RG4701 中有 Hyperbolic plane 的例子。有一個解微分方程的巧妙的方法。

2. RG3303 是 geodesic flows

3. RG5101 是 Jacobi 場一個 Jacobi field  $J(t)$  即沿著 geodesic  $\gamma$  滿足 Jacobi

equation 的向量場。  $\frac{D^2 J}{dt^2} + R(\gamma'(t), J(t))\gamma'(t) = 0$  這裡有關於測地線偏離的一點敘述。

4. 通過 geodesic 建構 exponential map 建立 normal coordinates 簡化計算。

5. Geodesic flows

$$\text{Geodesic equation : } \ddot{x}^i + \Gamma_{jk}^i \dot{x}^j \dot{x}^k = 0$$

Observe that any curve  $\gamma$  determines a curve  $t \rightarrow (\gamma(t), \gamma'(t))$  in  $TM$ .

If is a geodesic then in local coordinate  $t \rightarrow (x_1(t), \dots, x_n(t), \frac{dx_1(t)}{dt}, \dots, \frac{dx_n(t)}{dt})$  satisfies the first order system

$$(4) \quad \begin{cases} \frac{dx_k}{dt} = y_k \\ \frac{dy_k}{dt} = -\sum_{i,j} \Gamma_{ij}^k y_i y_j \end{cases}$$

on  $TU$ .

**Lemma 3.3.** *There exist a unique vector field  $G$  on  $TM$  whose trajectories are of the form  $t \rightarrow (\gamma(t), \gamma'(t))$ , where  $\gamma$  is a geodesic on  $M$ .*

**Definition 3.4.** *The vector field defined above is called the geodesic field on  $TM$  and its corresponding flow is called the geodesic flow on  $TM$ .*

6. [geodesic\(Wormhole\)](#)

§ 09 古典的作法如下：

$C: X = X(s)$  是曲面  $M$  上的一條曲線

$t = \frac{dX}{ds}$  是單位切向量， $\frac{dX}{ds} = X_i \frac{du^i}{ds}$  則

$$\frac{dt}{ds} = \frac{d}{ds} \left( X_i \frac{du^i}{ds} \right) = X_i \frac{d^2 u^i}{ds^2} + (X_{ij} \frac{du^j}{ds}) \frac{du^i}{ds}, \quad X_{ij} = \Gamma_{ij}^k X_k + b_{ij} N = \Gamma_{kj}^i X_k + b_{kj} N$$

$$= \left( \frac{d^2 u^i}{ds^2} + \Gamma_{jk}^i \frac{du^j}{ds} \frac{du^k}{ds} \right) X_i + b_{jk} N, \quad \text{注意到上式中 } i \leftrightarrow k, \text{ 且 } \Gamma_{ij}^k = \Gamma_{ji}^k$$

取切部，所以  $\frac{d^2 u^i}{ds^2} + \Gamma_{jk}^i \frac{du^j}{ds} \frac{du^k}{ds} = 0 \dots (1)$  是測地線方程式。

§ 10



杜武亮 p.108~110

有 Poincaré Half-Plane 上  
測地線的推導。

Fig. 14.4. Geodesics in the Poincaré half-plane.

### 11.1. Connection and curvature forms on the Poincaré disk

The *Poincaré disk* is the open unit disk

$$\mathbb{D} = \{z = x + iy \in \mathbb{C} \mid |z| < 1\}$$

in the complex plane with Riemannian metric

$$\langle \cdot, \cdot \rangle_z = \frac{4(dx \otimes dx + dy \otimes dy)}{(1 - |z|^2)^2} = \frac{4(dx \otimes dx + dy \otimes dy)}{(1 - x^2 - y^2)^2}.$$

An orthonormal frame for  $\mathbb{D}$  is

$$e_1 = \frac{1}{2}(1 - |z|^2) \frac{\partial}{\partial x}, \quad e_2 = \frac{1}{2}(1 - |z|^2) \frac{\partial}{\partial y}.$$

Find the connection matrix  $\omega = [\omega_j^i]$  and the curvature matrix  $\Omega = [\Omega_j^i]$  relative to the orthonormal frame  $e_1, e_2$  of the Riemannian connection  $\nabla$  on the Poincaré disk. (*Hint*: First find the dual frame  $\theta^1, \theta^2$ . Then solve for  $\omega_j^i$  in (11.9).)

Let  $D$  be the Poincare disk

1. Show that in polar coordinates  $(r, \theta)$ , the Poincare metric is given by

$$\langle \cdot, \cdot \rangle_{(r, \theta)} = \frac{4(dr^2 + r^2 d\theta^2)}{(1-r^2)^2}$$

Using polar coordinates to compute for  $D$

2. The Gaussian curvature
3. The Christoffel symbols
4. The geodesic equations

p.103~ 有 geodesic 古典性質

例如 *The speed of a geodesic on a Riemannian manifold is constant* ◦

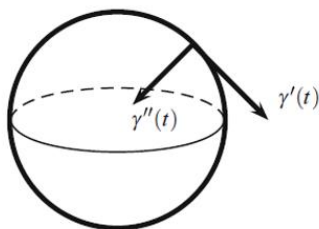
Let  $T = c'(t)$  be the velocity of the geodesic ◦

The speed is constant if and only if its square  $f(t) = \langle T, T \rangle$  is constant ◦

$$\text{But } f'(t) = \frac{d}{dt} \langle T, T \rangle = 2 \left\langle \frac{DT}{dt}, T \right\rangle = 0$$

So  $f(t)$  is constant ◦

## § 11 Geodesics on a sphere



Velocity and acceleration vectors of a great circle.

On a 2-sphere  $M$  of radius  $a$  in  $R^3$

Let  $\gamma(t)$  be a great circle parameterized by arc length ◦

Then  $\gamma(t)$  has unit speed ◦

Differentiating  $\langle \gamma'(t), \gamma'(t) \rangle = 1$  with respect

to  $t$  gives  $2 \langle \gamma''(t), \gamma'(t) \rangle = 0$

This shows that the acceleration  $\gamma''(t)$  of a unit-speed curve in  $R^3$  is perpendicular to the velocity ◦

Since  $\gamma(t)$  lies in the plane of the circle, so do  $\gamma'(t)$  and  $\gamma''(t)$  ◦

Being perpendicular to  $\gamma'(t)$ , the acceleration  $\gamma''(t)$  must in the radial

direction ◦ Hence, because  $\gamma''(t)$  is perpendicular to the tangent plane at  $\gamma(t)$  ◦



$$\frac{DT}{dt} = \left( \frac{dT}{dt} \right)_{\tan} = \gamma''(t)_{\tan} = 0$$

This shows that every great circle is a geodesic on the sphere ◦

[Differential Geometry in Physics] p.154 by Gabriel Lugo

Let  $S^2$  be a sphere of radius  $a$ , so that the metric is given by

$$ds^2 = a^2 d\theta^2 + a^2 \sin^2 \theta d\phi^2 \quad \text{and the Lagrangian is } L = a^2 \dot{\theta}^2 + a^2 \sin^2 \theta \dot{\phi}^2$$

The E-L equation for the  $\phi$  coordinate is  $\frac{d}{ds} \left( \frac{\partial L}{\partial \dot{\phi}} \right) - \frac{\partial L}{\partial \phi} = 0$

$$\frac{d}{ds} (2a^2 \sin^2 \theta \dot{\phi}) = 0 \quad \text{then} \quad \sin^2 \theta \frac{d\phi}{ds} = k$$

$$\sin^2 \theta d\phi = k ds$$

$$\sin^4 \theta d\phi^2 = k^2 ds^2 = k^2 (a^2 d\theta^2 + a^2 \sin^2 \theta d\phi^2)$$

$$(\sin^4 \theta - a^2 k^2 \sin^2 \theta) d\phi^2 = a^2 k^2 d\theta^2$$

$$\begin{aligned} d\phi &= \frac{akd\theta}{\sin \theta \sqrt{\sin^2 \theta - a^2 k^2}} = \frac{akd\theta}{\sin^2 \theta \sqrt{1 - a^2 k^2 \csc^2 \theta}} \\ &= \frac{ak \csc^2 \theta d\theta}{\sqrt{1 - a^2 k^2 (1 + \cot^2 \theta)}} = \frac{ak \csc^2 \theta d\theta}{\sqrt{(1 - a^2 k^2) - a^2 k^2 \cot^2 \theta}} \\ &= \frac{\csc^2 \theta d\theta}{\sqrt{\frac{1 - a^2 k^2}{a^2 k^2} - \cot^2 \theta}}, \quad \text{let } u = \cot \theta, \quad c^2 = \frac{1 - a^2 k^2}{a^2 k^2} \quad \text{then } du = -\csc^2 \theta d\theta \\ &= \frac{-du}{\sqrt{c^2 - u^2}} \quad \left( \frac{d}{dx} \sin^{-1} x = \frac{1}{\sqrt{1 - x^2}} \right) \end{aligned}$$

$$\phi = -\sin^{-1} \left( \frac{1}{c} \cot \theta \right) + \phi_0$$

$$\cot \theta = c \sin(\phi_0 - \phi)$$

$$\cos \theta = c \sin \theta (\sin \phi_0 \cos \phi - \cos \phi_0 \sin \phi)$$

$$a \cos \theta = (c \sin \phi_0)(a \sin \theta \cos \phi) - (c \cos \phi_0)(a \sin \theta \sin \phi)$$

$$z = Ax - By, \text{ where } A = c \sin \varphi_0, B = c \cos \varphi_0$$

We conclude that the geodesics of the sphere are great circles determined by the intersections with planes through the origin ◦

後面有旋轉面的 geodesics

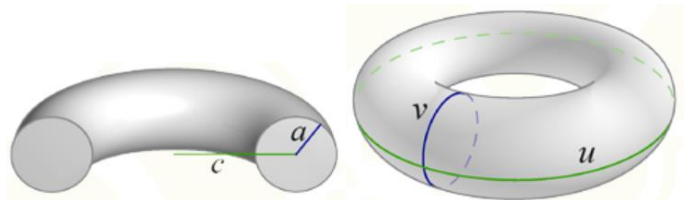
$$z=f(r) \text{ the first fundamental form is } ds^2 = (1+f'^2)dr^2 + r^2d\varphi^2$$

...

然後有一個有趣的蟲洞 Morris-Thorne Wormhole p.159

[geodesic\(Wormhole\)](#)

§ 12 geodesics on a torus



[[the torus](#)]

$$\begin{cases} x = c + a \cos v \cos u \\ y = c + a \cos v \sin u \\ z = a \sin v \end{cases}$$

By geodesic equations  $\ddot{x}^i + \Gamma_{jk}^i \dot{x}^j \dot{x}^k = 0$  we get

$$\ddot{u} - \frac{2a \sin v}{c + a \cos v} uv = 0$$

$$\ddot{v} + \frac{1}{a} \sin v c + a \cos v u^2 = 0$$

Using the substitution  $w=c+a\cos v$  and integrating (with a trick or two) gives a solution in terms of  $u$  and  $v$

$$\begin{cases} \dot{u} = \frac{k}{c + a \cos v^2} \\ \dot{v} = \pm \sqrt{-\frac{k^2}{a^2 c + a \cos v^2} + l} \end{cases}$$

The Clairaut parameterization of a torus treats it as a surface of revolution ◦

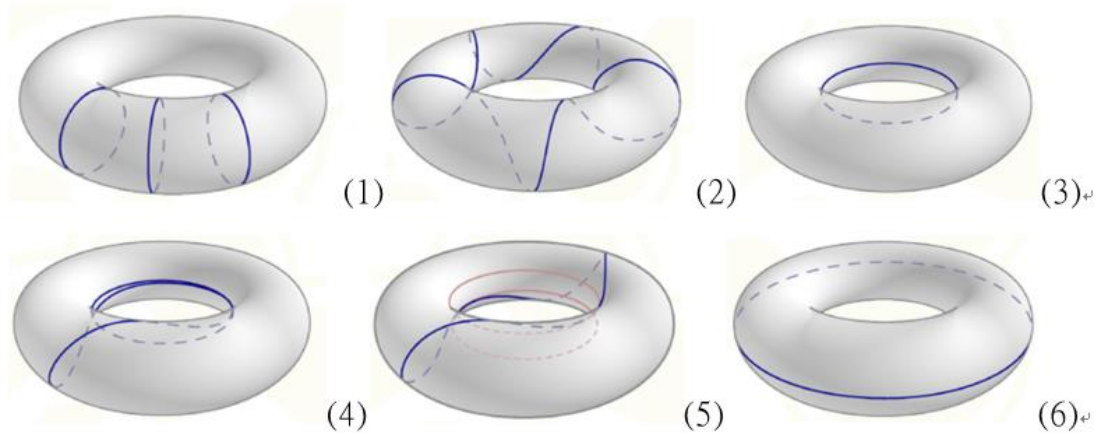
From it, we obtain a formula for  $\frac{du}{dv}$

$$\frac{du}{dv} = \pm \frac{ah}{c + a \cos v \sqrt{c + a \cos v^2 - h^2}}$$

$$h = c + a \cos v \sin \varphi$$

The five families of geodesics :

- (1)  $h=0$  , these are the meridians
- (2)  $0 < h < c-a$  unbounded geodesics , which alternately cross inner and outer equators ◦
- (3)  $h=c-a$  the inner equator , and geodesics asymptotic to it
- (4)  $c-a < h < c+a$  bounded geodesics, which cross the outer equator but bounce off barrier curves
- (5)  $h=c+a$  the outer equator



When  $h > c+a$  , there are no real solutions to the formula for  $\frac{du}{dv}$

### § 13 The geodesic equations for a metric of a BH

We consider a charged BH surrounded quintessence with the equation of state

parameter  $\varepsilon = \frac{p\phi}{\sigma\phi}$  ,  $-1 < \varepsilon < -\frac{1}{3}$  , and  $\alpha$  is the normalization constant ◦

Quintessence 第五元素 a hypothetical form of dark energy , a scalar field

The metric of a charged BH reads :

$$ds^2 = f(r)dt^2 - \frac{1}{f(r)}dr^2 - r^2(d\theta^2 + \sin^2\theta d\phi^2)$$

$$\text{Where } f(r) = 1 - \frac{2M}{r} + \frac{Q^2}{r} - \frac{\alpha}{r^{3\varepsilon+1}}$$

Where M and Q represent the mass and charge of the BH respectively ◦

The metric represents a BH for  $M > Q$  , an extremal BH for  $M=Q$  and a naked singularity

for  $M < Q$  .

The geodesic equations for the metric are given by

$$\ddot{t} + \frac{f'(r)}{f(r)} \dot{r} \dot{t} = 0 \quad (1)$$

$$\ddot{r} + \left[ \frac{f'(r)\dot{t}^2 + f^{-1}(r)'r^2 - 2r\dot{\theta}^2 - 2r\sin^2\theta\dot{\phi}^2}{2f^{-1}(r)} \right] = 0 \quad (2)$$

$$\ddot{\theta} + \frac{2}{r} \dot{r} \dot{\theta} - \cos\theta \sin\theta \dot{\phi}^2 = 0 \quad (3)$$

$$\ddot{\phi} + \frac{2}{r} \dot{r} \dot{\phi} + 2 \cot\theta \dot{\theta} \dot{\phi} = 0 \quad (4)$$

Where the prime denotes the differentiation with respect to  $r$  .

The geodesic equations are  $\ddot{x}^i + \Gamma_{jk}^i \dot{x}^j \dot{x}^k = 0$

Where  $\Gamma_{jk}^i = \frac{1}{2} g^{il} \left( \frac{\partial g_{kl}}{\partial x^j} + \frac{\partial g_{jl}}{\partial x^k} - \frac{\partial g_{jk}}{\partial x^l} \right)$

$$(g_{ij}) = \begin{pmatrix} f(r) & 0 & 0 & 0 \\ 0 & -\frac{1}{f(r)} & 0 & 0 \\ 0 & 0 & -r^2 & 0 \\ 0 & 0 & 0 & -r^2 \sin^2 \theta \end{pmatrix}, \quad (g^{ij}) = \begin{pmatrix} \frac{1}{f(r)} & 0 & 0 & 0 \\ 0 & -f(r) & 0 & 0 \\ 0 & 0 & -\frac{1}{r^2} & 0 \\ 0 & 0 & 0 & \frac{-1}{r^2 \sin^2 \theta} \end{pmatrix}$$

$$X = X(t, r, \theta, \phi)$$

$$\ddot{t} + \Gamma_{jk}^1 \dot{x}^j \dot{x}^k = 0$$

$$\ddot{t} + \Gamma_{11}^1 \dot{t} \dot{t} + \Gamma_{12}^1 \dot{t} \dot{r} + \Gamma_{21}^1 \dot{r} \dot{t} + \Gamma_{22}^1 \dot{r} \dot{r} + \{ \dots \}$$

$$\Gamma_{11}^1 = \frac{1}{2} g^{1l} \left( \frac{\partial g_{1l}}{\partial x^1} + \frac{\partial g_{1l}}{\partial x^1} - \frac{\partial g_{11}}{\partial x^l} \right) = 0 ,$$

$$\Gamma_{22}^1 = \frac{1}{2} g^{1l} \left( \frac{\partial g_{2l}}{\partial x^2} + \frac{\partial g_{2l}}{\partial x^2} - \frac{\partial g_{22}}{\partial x^l} \right) = \frac{1}{2} g^{11} \left( \frac{\partial g_{21}}{\partial x^2} + \frac{\partial g_{21}}{\partial x^2} - \frac{\partial g_{22}}{\partial x^1} \right) = 0$$

$$\Gamma_{12}^1 = \frac{1}{2} g^{1l} \left( \frac{\partial g_{2l}}{\partial x^1} + \frac{\partial g_{1l}}{\partial x^2} - \frac{\partial g_{12}}{\partial x^l} \right) = \frac{1}{2} \times \frac{1}{f(r)} \times \frac{\partial f}{\partial r} = \Gamma_{21}^1$$

所以得到(1)  $\ddot{t} + \frac{f'(r)}{f(r)} \dot{r} \dot{t} = 0$

$$\ddot{r} + \Gamma_{jk}^2 \dot{x}^j \dot{x}^k = 0$$

$$\Gamma_{11}^2 = \frac{1}{2} g^{2l} \left( \frac{\partial g_{1l}}{\partial x^1} + \frac{\partial g_{1l}}{\partial x^1} - \frac{\partial g_{11}}{\partial x^l} \right) = \frac{1}{2} g^{22} \times \left( -\frac{\partial g_{11}}{\partial x^2} \right) = \frac{1}{2} f(r) \frac{\partial f(r)}{\partial r}$$

$$\Gamma_{22}^2 = \frac{1}{2} g^{22} \left( \frac{\partial g_{2l}}{\partial x^2} + \frac{\partial g_{2l}}{\partial x^2} - \frac{\partial g_{22}}{\partial x^l} \right) = \frac{1}{2} g^{22} \left( \frac{\partial g_{22}}{\partial r} + \frac{\partial g_{22}}{\partial r} - \frac{\partial g_{22}}{\partial r} \right) = \frac{1}{2} f(r) \times f^{-1}(r)'$$

$$\Gamma_{33}^2 = \frac{1}{2} g^{2l} \left( \frac{\partial g_{3l}}{\partial x^3} + \frac{\partial g_{3l}}{\partial x^3} - \frac{\partial g_{33}}{\partial x^l} \right) = \frac{1}{2} (-2rf(r))$$

$$\Gamma_{44}^2 = \frac{1}{2} g^{2l} \left( \frac{\partial g_{4l}}{\partial x^4} + \frac{\partial g_{4l}}{\partial x^4} - \frac{\partial g_{44}}{\partial x^l} \right) = \frac{1}{2} (-f(r))(2r \sin^2 \theta)$$

得到(2)

$$\ddot{\theta} + \Gamma_{jk}^3 \dot{x}^j \dot{x}^k = 0$$

$$\Gamma_{11}^3 = 0 \quad \Gamma_{22}^3 = 0 \quad \Gamma_{33}^3 = 0 \quad \Gamma_{44}^3 = -\sin \theta \cos \theta$$

$$\Gamma_{23}^3 = \Gamma_{32}^3 = \frac{1}{r}, \text{ 其他皆為 } 0$$

$$\text{得到(3)} \quad \ddot{\theta} + \frac{2}{r} \dot{r} \dot{\theta} - \cos \theta \sin \theta \dot{\phi}^2 = 0$$

$$\ddot{\phi} + \Gamma_{jk}^4 \dot{x}^j \dot{x}^k = 0$$

$$\Gamma_{24}^4 = \Gamma_{42}^4 = \frac{1}{r} \quad \Gamma_{34}^4 = \Gamma_{43}^4 = \cot \theta, \text{ 其他皆為 } 0$$

$$\text{得到(4)} \quad \ddot{\phi} + \frac{2}{r} \dot{r} \dot{\phi} + 2 \cot \theta \dot{\theta} \dot{\phi} = 0$$

## § 14 參考書目

1. An Introduction to Riemannian Geometry by Jose Natario
2. 大域微分幾何 黃武雄
3. Spacetime and Geometry by Sean Carroll
4. Riemannian geometry and geometric Analysis by Jurgen Jost p.23
5. Differential Geometry in Physics by Gabriel Lugo p.154