

An Introduction to Riemannian Geometry CH 6.5 p.272

Spacetime and Geometry CH5.1 p.193

By Birkhoff Theorem :

In GR , the unique spherically symmetric vacuum solution is the Schwarzschild metric ◦

$$ds^2 = -\left(1 - \frac{2GM}{r}\right)dt^2 + \left(1 - \frac{2GM}{r}\right)^{-1}dr^2 + r^2 d\Omega^2$$

Where $d\Omega^2 = d\theta^2 + \sin^2 \theta d\phi^2$

Constant M is the mass of the gravitating object ◦

Since we are interested in the solution outside a spherical body , we care about Einstein

equation in vacuum , $R_{\mu\nu} = 0$

§ The Schwarzschild Solution

$$g = -A^2(r)dt^2 + B^2(r)dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2$$

For unknown positive smooth function A , B : $R \rightarrow R$

$$\omega^0 = A(r)dt, \omega^1 = B(r)dr, \omega^2 = r d\theta, \omega^3 = r \sin \theta d\phi$$

$\{\omega^0, \omega^1, \omega^2, \omega^3\}$ is an orthonormal coframe ◦

$$d\omega^i = \sum_{j=0}^3 \omega^j \wedge \omega_j^i$$

$\omega_0^0 = \omega_i^i = 0, \omega_i^0 = \omega_0^i, \omega_j^i = -\omega_i^j$ on an orthonormal frame ◦

$$d\omega^0 = A'(r)dr \wedge dt = \frac{A'}{B} \omega^1 \wedge dt$$

$$= \omega^0 \wedge \omega_0^0 + \omega^1 \wedge \omega_1^0 + \omega^2 \wedge \omega_2^0 + \omega^3 \wedge \omega_3^0$$

$$\therefore \omega_1^0 = \frac{A'}{B} dt, \omega_2^0 = \omega_3^0 = 0$$

$$d\omega^1 = 0$$

$$d\omega^2 = dr \wedge d\theta = \frac{1}{B} \omega^1 \wedge d\theta$$

$$d\omega^3 = \sin \theta dr \wedge d\phi + r \cos \theta d\theta \wedge d\phi = \frac{\sin \theta}{B} \omega^1 \wedge d\phi + \cos \theta \omega^2 \wedge d\phi$$

$$= \omega^0 \wedge \omega_0^3 + \omega^1 \wedge \omega_1^3 + \omega^2 \wedge \omega_2^3 + \omega^3 \wedge \omega_3^3$$

$$\therefore \omega_1^3 = \frac{\sin \theta}{B} d\phi$$

Yield the nonvanishing connection forms :

$$\omega_1^0 = \omega_0^1 = \frac{A'}{B} dt ; \omega_1^2 = -\omega_2^1 = \frac{1}{B} d\theta ; \omega_1^3 = -\omega_3^1 = \frac{\sin \theta}{B} d\varphi$$

$$\omega_2^3 = -\omega_3^2 = \cos \theta d\varphi$$

$$\Omega_i^j = d\omega_i^j - \omega_i^k \wedge \omega_k^j = \sum_{m < n} R_{nm}^j \omega^m \wedge \omega^n$$

Then the curvature forms :

$$\Omega_r^0 = \Omega_0^r = \frac{A''B - A'B'}{AB^3} \omega^r \wedge \omega^0;$$

$$\Omega_\theta^0 = \Omega_0^\theta = \frac{A'}{rAB^2} \omega^\theta \wedge \omega^0;$$

$$\Omega_\varphi^0 = \Omega_0^\varphi = \frac{A'}{rAB^2} \omega^\varphi \wedge \omega^0;$$

$$\Omega_r^\theta = -\Omega_\theta^r = \frac{B'}{rB^3} \omega^\theta \wedge \omega^r;$$

$$\Omega_r^\varphi = -\Omega_\varphi^r = \frac{B'}{rB^3} \omega^\varphi \wedge \omega^r;$$

$$\Omega_\theta^\varphi = -\Omega_\varphi^\theta = \frac{B^2 - 1}{r^2 B^2} \omega^\varphi \wedge \omega^\theta.$$

例如 $\Omega_r^\theta = \Omega_1^2 = d\omega_1^2 - \{\omega_1^0 \wedge \omega_0^2 + \omega_1^1 \wedge \omega_1^2 + \omega_1^2 \wedge \omega_2^2 + \omega_1^3 \wedge \omega_3^2\}$

$$= \frac{-B'}{B^2} dr \wedge d\theta = \frac{-B}{rB^3} \omega^1 \wedge \omega^2 \quad \text{as } \omega^1 = Bdr, \omega^2 = rd\theta$$

$$\Omega_\theta^\varphi = \Omega_2^3 = d\omega_2^3 - (\omega_2^0 \wedge \omega_0^3 + \omega_2^1 \wedge \omega_1^3 + \omega_2^2 \wedge \omega_2^3 + \omega_2^3 \wedge \omega_3^3)$$

$$= d(\cos \theta d\varphi) - \left(-\frac{1}{B} d\theta \wedge \frac{\sin \theta}{B} d\varphi\right) = \frac{-1}{r^2} \omega^2 \wedge \omega^3 + \frac{1}{r^2 B^2} \omega^2 \wedge \omega^3 = \frac{-B^2 + 1}{r^2 B^2} \omega^2 \wedge \omega^3$$

Ricci tensor

$$R_{ij}^j = \Omega_i^j(E_i, E_j) \quad R_{ij} = \sum_k R_{kij}^k$$

$$R_{00} = R_{000}^0 + R_{100}^1 + R_{200}^2 + R_{300}^3 = -(R_{010}^1 + R_{020}^2 + R_{030}^3)$$

$$= \frac{A''B - A'B'}{AB^3} + \frac{2A'}{rAB^2}$$

Where

$$R_{010}^1 = \Omega_0^1(E_0, E_1) = -\frac{A''B - A'B'}{AB^3}, \quad R_{020}^2 = \Omega_0^2(E_0, E_\theta) = -\frac{A'}{rAB^2}$$

$$R_{030}^3 = \Omega_0^\varphi(E_0, E_\varphi) = \frac{-A'}{rAB^2}$$

$$R_{11} = R_{011}^0 + R_{111}^1 + R_{211}^2 + R_{311}^3 = -(R_{101}^0 + R_{121}^2 + R_{131}^3) = \frac{-A''B + A'B'}{AB^3} + \frac{2B'}{rB^3}$$

Where

$$R_{101}^0 = \Omega_1^0(E^1, E^0) = \frac{-A''B + A'B'}{AB^3}, \quad R_{121}^2 = \Omega_1^2(E_1, E_2) = \frac{-B'}{rB^3},$$

$$R_{131}^3 = \Omega_1^3(E_1, E_3) = \frac{-B'}{rB^3}$$

$$R_{22} = R_{022}^0 + R_{122}^1 + R_{222}^2 + R_{322}^3 = -R_{202}^0 - R_{212}^1 - R_{232}^3$$

$$= -\frac{A'}{rAB} - \frac{B'}{r^3B} - \frac{B^2-1}{r^2B}$$

Where

$$R_{202}^0 = \Omega_\theta^0(E_\theta, E_0) = \frac{A'}{rAB^2}, \quad R_{212}^1 = \Omega_\theta^1(E_\theta, E_r) = \frac{-B'}{rB^3}, \quad R_{232}^3 = \Omega_\theta^3(E_\theta, E_\varphi) = -\frac{B^2-1}{r^2B^2}$$

$$R_{33} = R_{033}^0 + R_{133}^1 + R_{233}^2 + R_{333}^3 = R_{22}$$

$$R_{00} = \frac{A''B - A'B'}{AB^3} + \frac{2A'}{rAB^2};$$

$$R_{rr} = -\frac{A''B - A'B'}{AB^3} + \frac{2B'}{rB^3};$$

$$R_{\theta\theta} = R_{\varphi\varphi} = -\frac{A'}{rAB^2} + \frac{B'}{rB^3} + \frac{B^2-1}{r^2B^2}.$$

Thus the vacuum Einstein field equation $Ric = 0$ is equivalent to the ODE system

$$\begin{cases} \frac{A''}{A} - \frac{A'B'}{AB} + \frac{2A'}{rA} = 0 \\ \frac{A''}{A} - \frac{A'B'}{AB} - \frac{2B'}{rB} = 0 \\ \frac{A'}{A} - \frac{B'}{B} - \frac{B^2-1}{r} = 0 \end{cases} \Leftrightarrow \begin{cases} \frac{A'}{A} + \frac{B'}{B} = 0 \\ \left(\frac{A'}{A}\right)' + 2\left(\frac{A'}{A}\right)^2 + \frac{2A'}{rA} = 0 \\ \frac{2B'}{B} + \frac{B^2-1}{r} = 0 \end{cases} .$$

The last equation can be immediately solved to yield $B = \left(1 - \frac{2m}{r}\right)^{-\frac{1}{2}}$, where $m \in \mathbb{R}$ is an integration constant ◦

The first equation implies that $A = \frac{\alpha}{B}$ for some constant $\alpha > 0$ ◦

By rescaling the time coordinate t we can assume that $\alpha = 1$

Finally, it is easy to check that the second ODE is identically satisfied ◦

Therefore there exists a one-parameter family of solution of vacuum Einstein field equation of the form we sought, given by

$$g = -\left(1 - \frac{2m}{r}\right)dt^2 + \left(1 - \frac{2m}{r}\right)^{-1}dr^2 + r^2d\theta^2 + r^2\sin^2\theta d\varphi^2$$

§ 5.6 Schwarzschild black hole [Spacetime and Geometry] p.218